

LANDFORM IDENTIFICATION - LUNAR RADAR IMAGES

Henry J. Moore, U.S. Geological Survey, Menlo Park, CA 94025
T. W. Thompson, Jet Propulsion Laboratory, Pasadena, CA 91109

Three sets of polarized radar-echo images of the Moon are being examined to establish the relation between radar resolution and landform-identification resolution [1-4]. The wavelengths, radar resolutions (cell sizes), and approximate number of real or apparent landforms for the sets are as follows:

<u>Set #</u>	<u>Wavelength</u>	<u>Cell sizes</u>	<u>Number of landforms</u>
1	3.8 cm [5]	1-2 km	1,553
2	70 cm [6] (high resolution)	2.5-5 km	1,594
3	70 cm [7] (low resolution)	10-20 km	983

The results of the study should be valuable to those planning to acquire or interpret radar images of the Earth or other planetary bodies.

After comparison with lunar maps and photographs, real and apparent landforms on the radar images are grouped into one of seven classes [1-4]: (1) resolved and clearly identified; (2) resolved and would probably be correctly identified; (3) resolved, but interpretation is uncertain; (4) detected, but elements are not resolved; (5) not detected; (6) array of landforms is resolved, but interpretation of the array is uncertain; and (7) radar portrays a fictitious landform.

Data recorded for each real or apparent landform for each set of images includes the following: (1) a name, (2) selenographic coordinates, (3) diameter and relief obtained from lunar maps and photographs, (4) the class, (5) diameter measured on the radar image, (6) background terrain, (7) geologic age, and (8) the Lunar Aeronautical Chart number. A computer program sorts and orders the data by diameter or relief, computes the percent of each class in frequency bins of 100, and computes the geometric mean of the landform diameter or relief for each frequency bin. Calculations are made in frequency steps of 10.

Current results show strong relations between radar resolution and diameter or relief of landforms that are

clearly identified and those that would probably be correctly identified (class 1 + class 2), as shown in table 1. Current results are not depicted; they are similar, but not identical, the those in previous abstracts [e.g. 1, figure 1].

Table 1. Percentage of resolved and identified landforms portrayed in lunar radar images.

Percentage of class 1+2 landforms	Mean diameters of landforms (km) corresponding to the indicated percentage.		
	3.8 cm	70 cm high	70 cm low
10	5	12	35
20	6	16	56
30	7	25	--
40	9	35	--
50	11	44	--
60	16	--	--
70	20	--	--
80	35	--	--

Percentage of class 1+2 landforms	Mean relief of landforms (km) corresponding to the indicated percentage.		
	3.8 cm	70 cm high	70 cm low
10	0.5	1.0	2.7
20	0.6	1.7	---
30	0.8	2.6	---
40	1.1	2.8	---
50	1.4	3.1	---
60	2.0	---	---
70	2.2	---	---
80	2.6	---	---

Landforms are not detected (class 5) at all diameters and reliefs, but the percentage of undetected landforms decreases with increasing mean diameter and increasing mean relief. Landforms are simply detected (class 4) at most mean diameters and reliefs. Ambiguous arrays (class 6) portrayed by the radar constitute up to about 16, 22, and 15 percent of the landforms at various diameters and relief values for the 3.8 cm, 70 cm high resolution, and 70 cm low resolution images, respectively. Only a few percent of the landforms portrayed by the radar images at various diameters and relief values are fictitious (class 7).

When data acquisition is complete, the data will be analyzed as functions of angle of incidence (lunar-scattering function) [6], background terrain, and geologic age [7].

Preliminary comparisons of the actual observed crater

size-frequency distributions with those obtained from the radar images show increasing departures with increasing resolution. For the 3.8 cm radar images, the cumulative frequency of craters greater than 22.6 km across agree to within 21 percent of the actual cumulative frequency, and the population indices (slope of the distribution) are similar and near -2. Here, the crater frequencies from the radar image lie below the observed ones. Comparisons for the other radar images are less satisfactory at this time.

References

- [1] Moore, H. J. and Thompson, T. W., 1986, Landform identification in lunar radar images: *Lunar and Planet. Science XVII*, p. 565-566.
- [2] Moore, H. J. and Thompson, T. W., 1986, Lunar radar images - landform identification: *Repts. Planet. Geol. and Geophys. Prog.*--1985, NASA TM 88383, p. 550-552.
- [3] Thompson, T. W. and Moore, H. J., 1985, Landform identification in radar images: *Lunar and Planet. Science XVI*, p. 860-861.
- [4] Moore, H. J. and Thompson, T. W., 1985, Landform identification on radar images: *Repts. Planet. Geol. and Geophys. Prog.*--1984, NASA TM 87563, p. 458-460.
- [5] Zisk, S. H., Pettengill, G. W., and Catuna, G. W., 1974, High-resolution radar maps of the lunar surface at 3.8-cm wavelength: *The Moon*, v.10, p. 17-50.
- [6] Thompson, T. W., unpublished data.
- [7] Thompson, T. W., 1974, Atlas of lunar radar maps at 70 cm wavelength: *The Moon*, v.10, p. 51-85.
- [8] Hagfors, T., 1970, Remote probing of the Moon by infrared and microwave emissions and by radar: *Radio Science*, v. 5, p. 189-227.
- [9] Wilhelms, D. E., 1980, Stratigraphgy of part of the lunar near side, U. S. Geol. Survey Prof. Paper 1046-A, 71p.